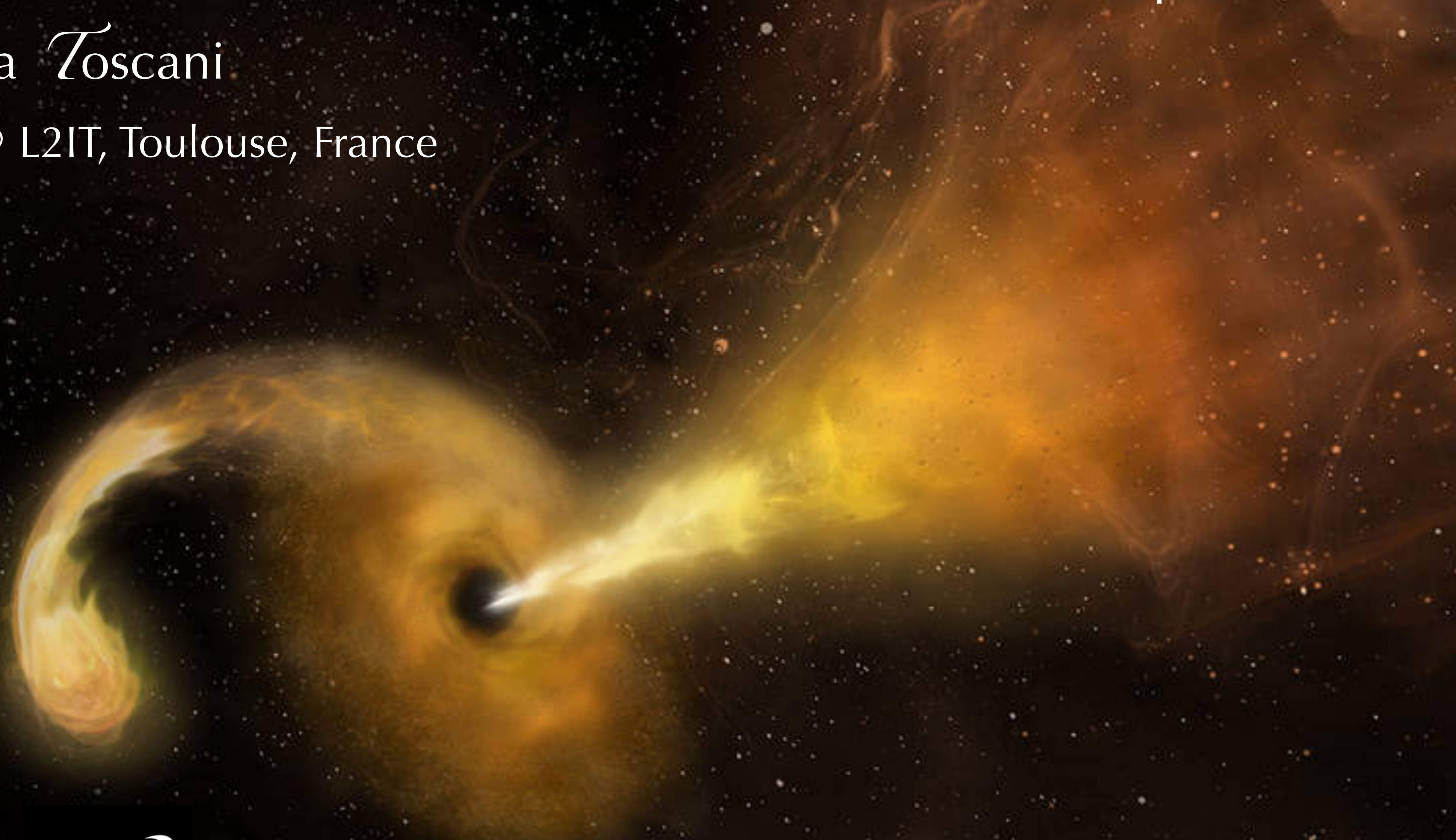


Gravitational waves with PHANTOM: the case of *Tidal Disruption Events*

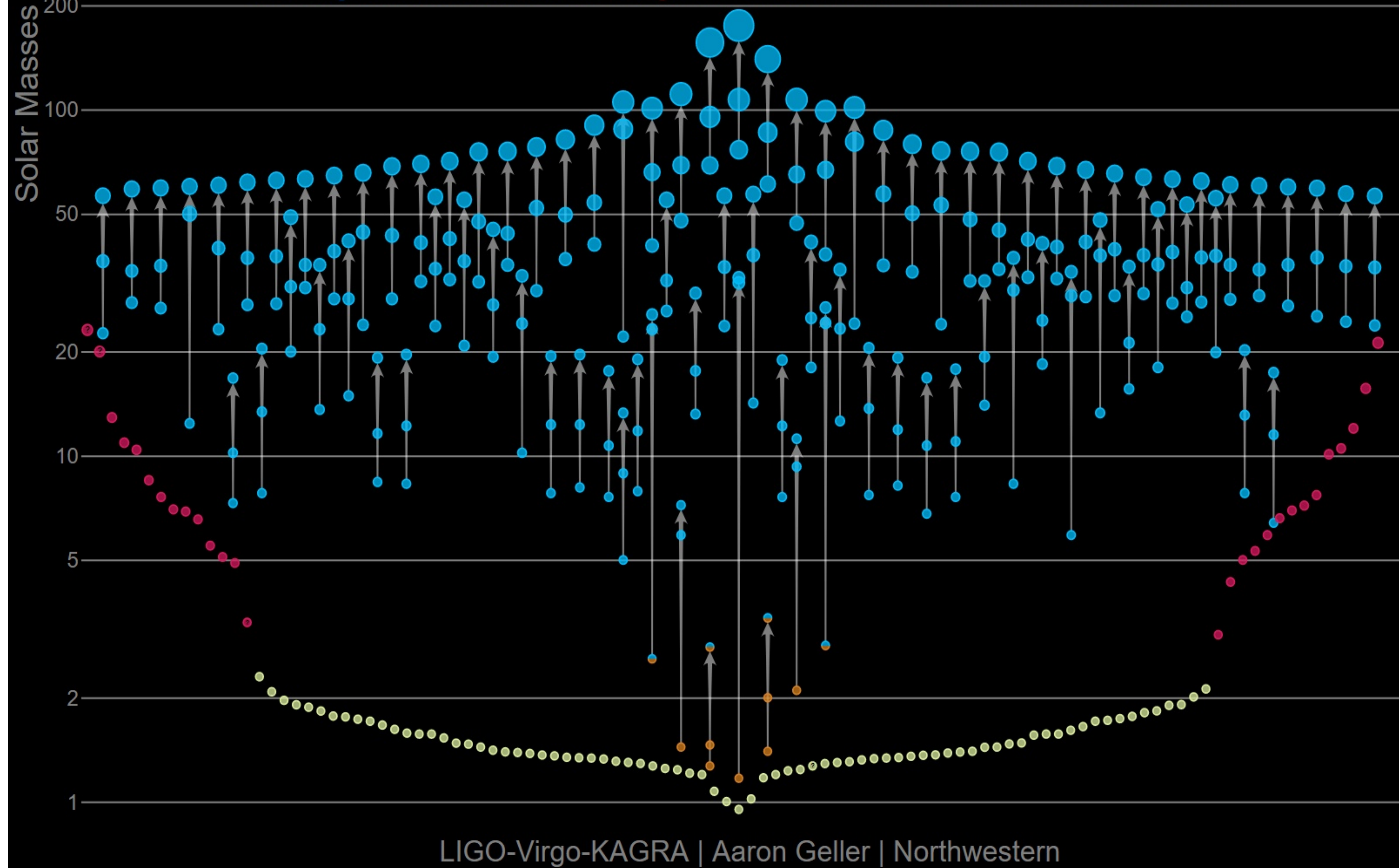
Martina Toscani

Postdoc @ L2IT, Toulouse, France



Masses in the Stellar Graveyard

LIGO-Virgo-KAGRA Black Holes LIGO-Virgo-KAGRA Neutron Stars EM Black Holes EM Neutron Stars



Since 2015,
90ish gravitational wave
(GW) events detected!

Revolution in astronomy

LIGO-Virgo-KAGRA | Aaron Geller | Northwestern

Ground based laser interferometers

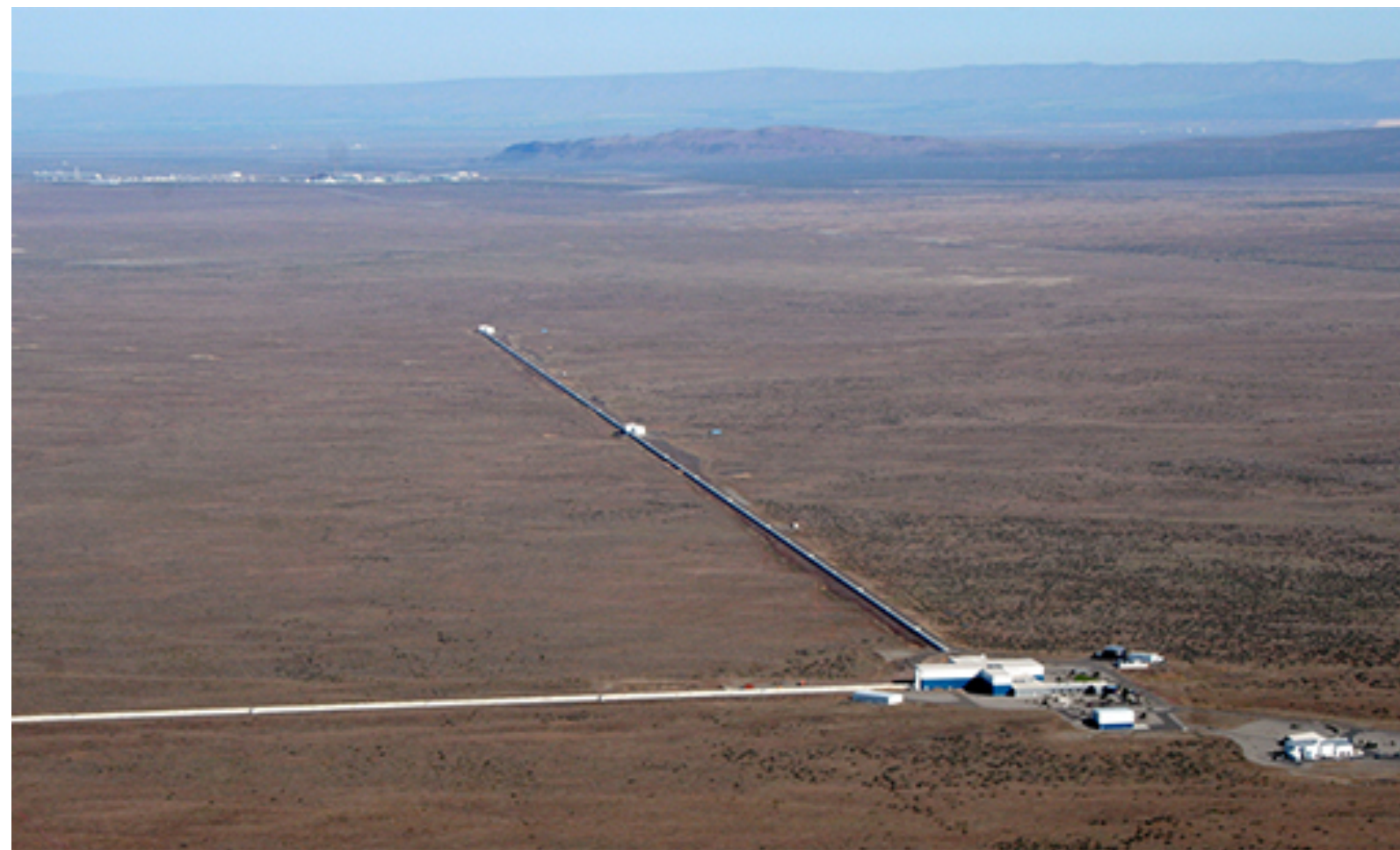
Most sensitive instrument on Earth
1000 times smaller than an atom

10 ~1000 Hz

LIGO | Handford

LIGO | Livingston

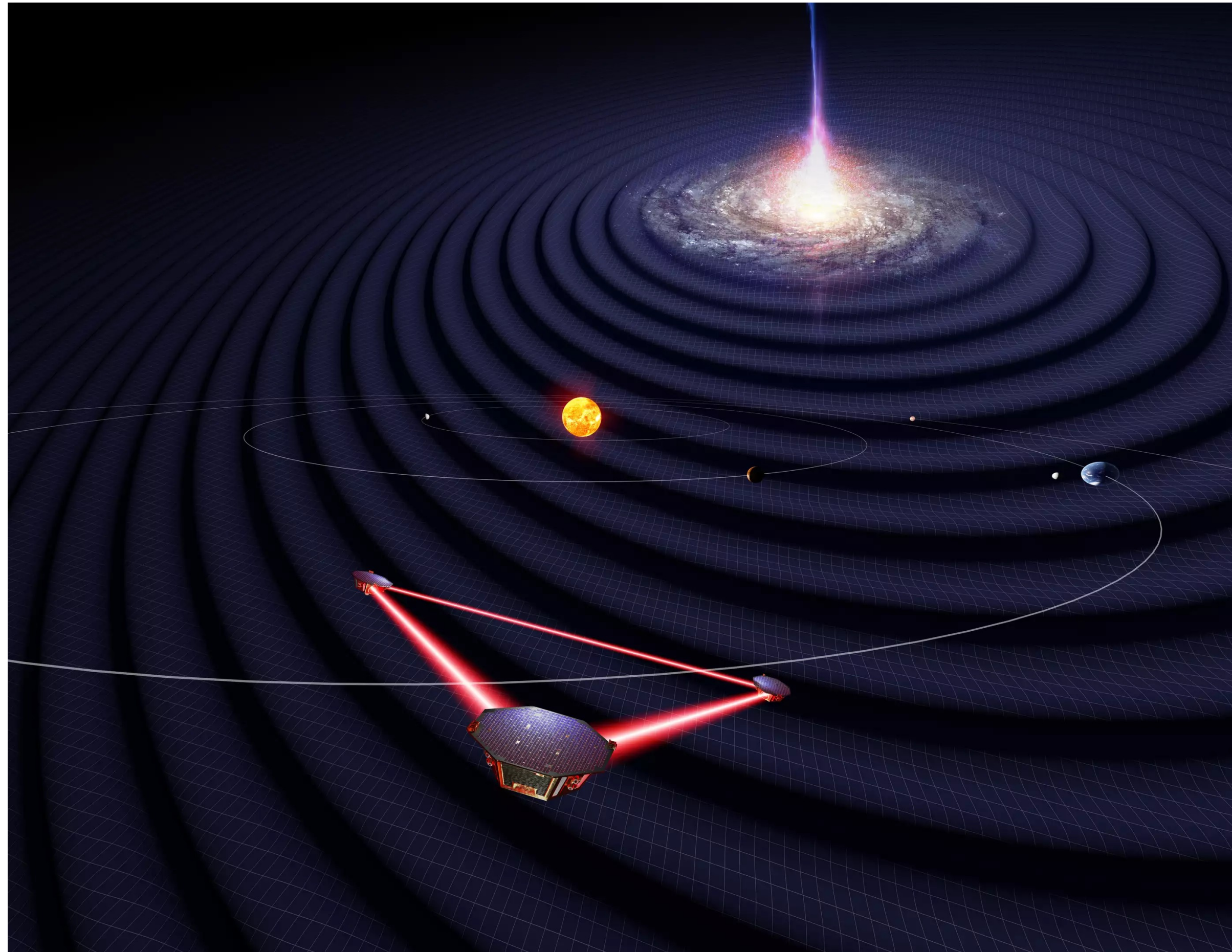
Virgo



+ KAGRA

Credits: LIGO Laboratory (first two images) and Virgo / Nicola Baldocchi 2015

Space based laser interferometers

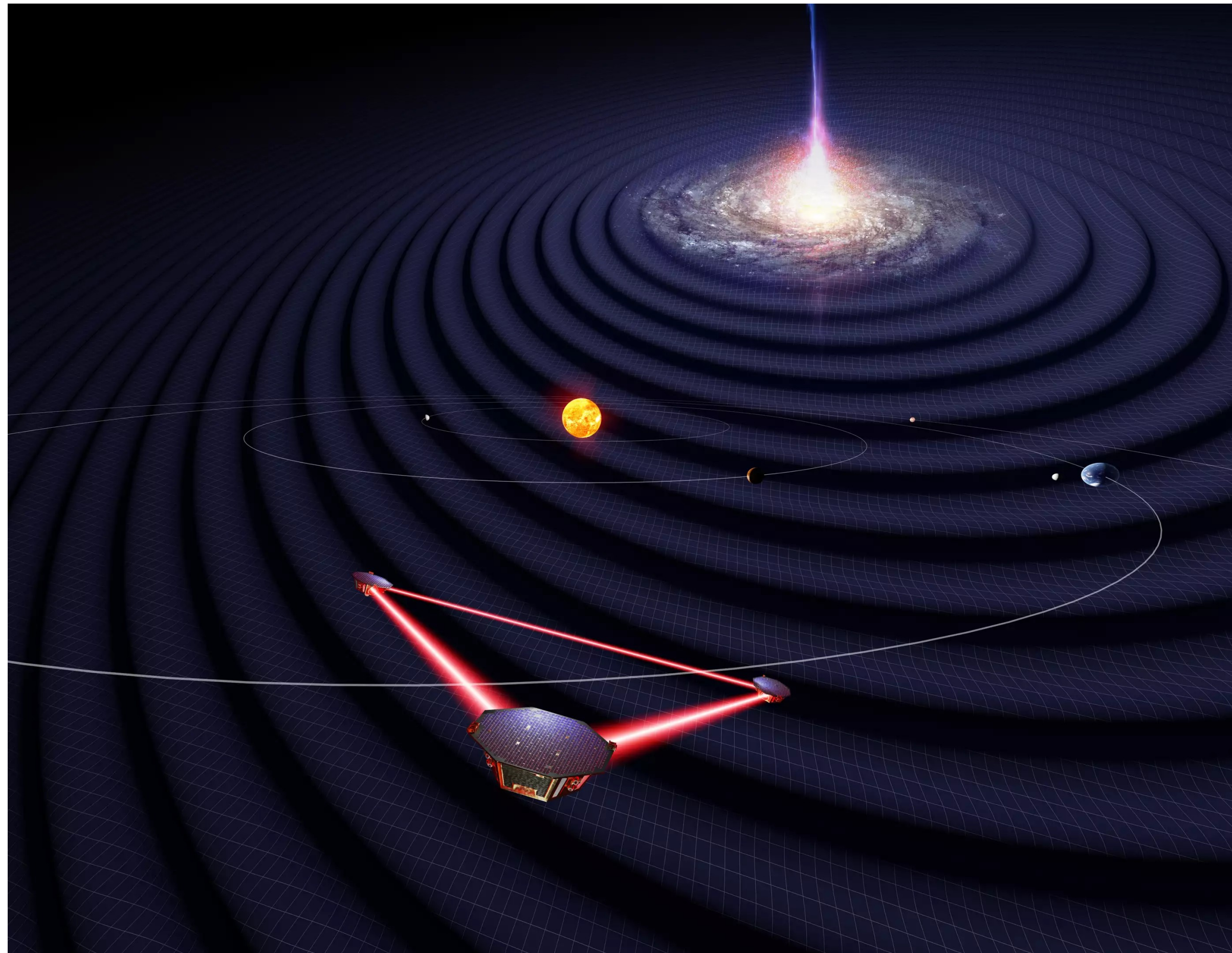


Credits: University of Florida / Simon Barke (CC BY 4.0)

LISA is the future! ~2037 (Amaro-Seoane+ 2017, 2022)

0.0001 ~ 0.1 Hz

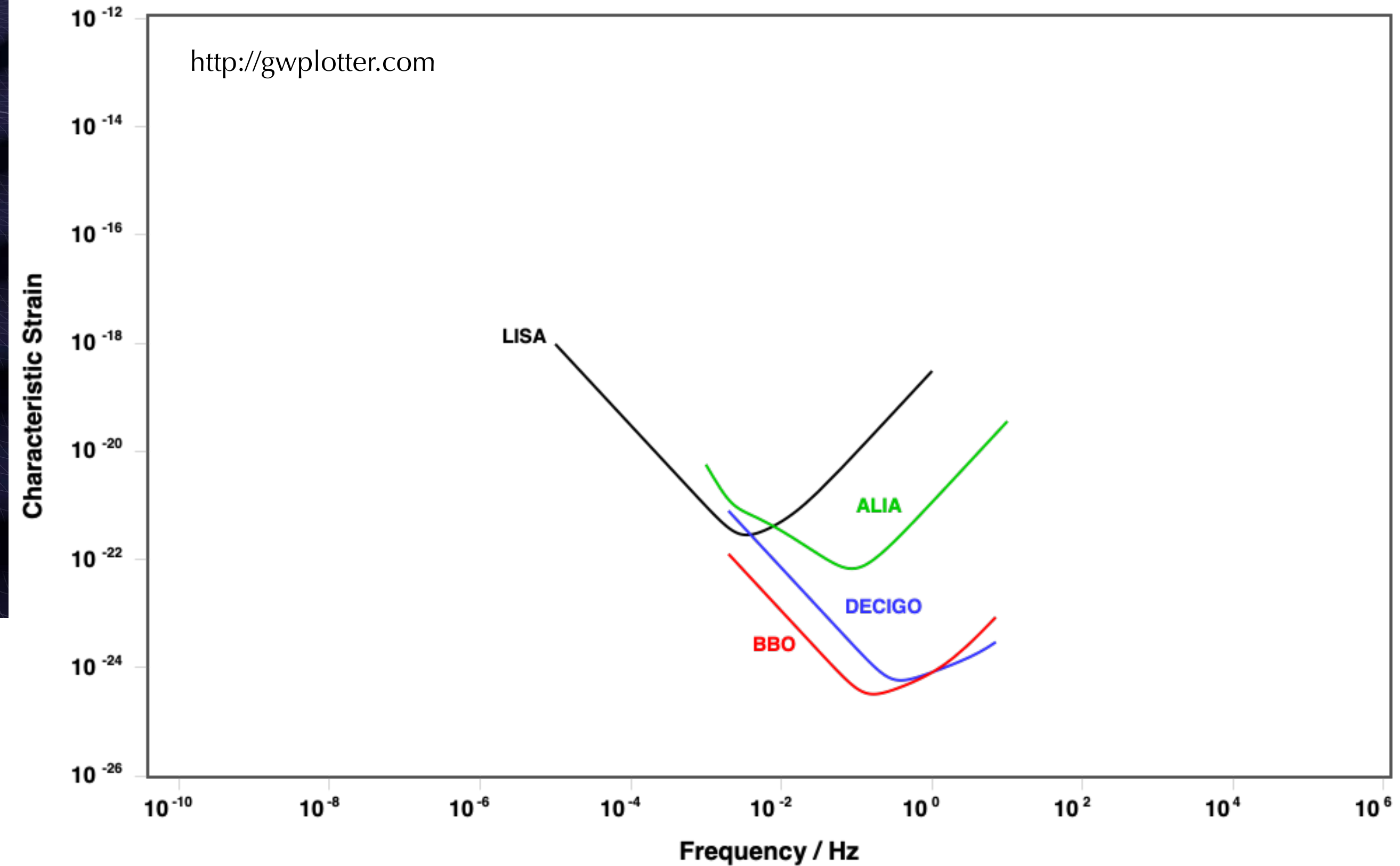
Space based laser interferometers



Credits: University of Florida / Simon Barke (CC BY 4.0)

LISA is the future! ~2037 (Amaro-Seoane+ 2017, 2022)

0.0001 ~ 0.1 Hz



General Relativity in a nutshell

Einstein field equations

$$\text{curvature } G_{\mu\nu} \propto T_{\mu\nu} \text{ mass and energy}$$

General Relativity in a nutshell

Einstein field equations

curvature $G_{\mu\nu} \propto T_{\mu\nu}$ mass and energy

Accelerated, asymmetric mass distribution
perturbs spacetime

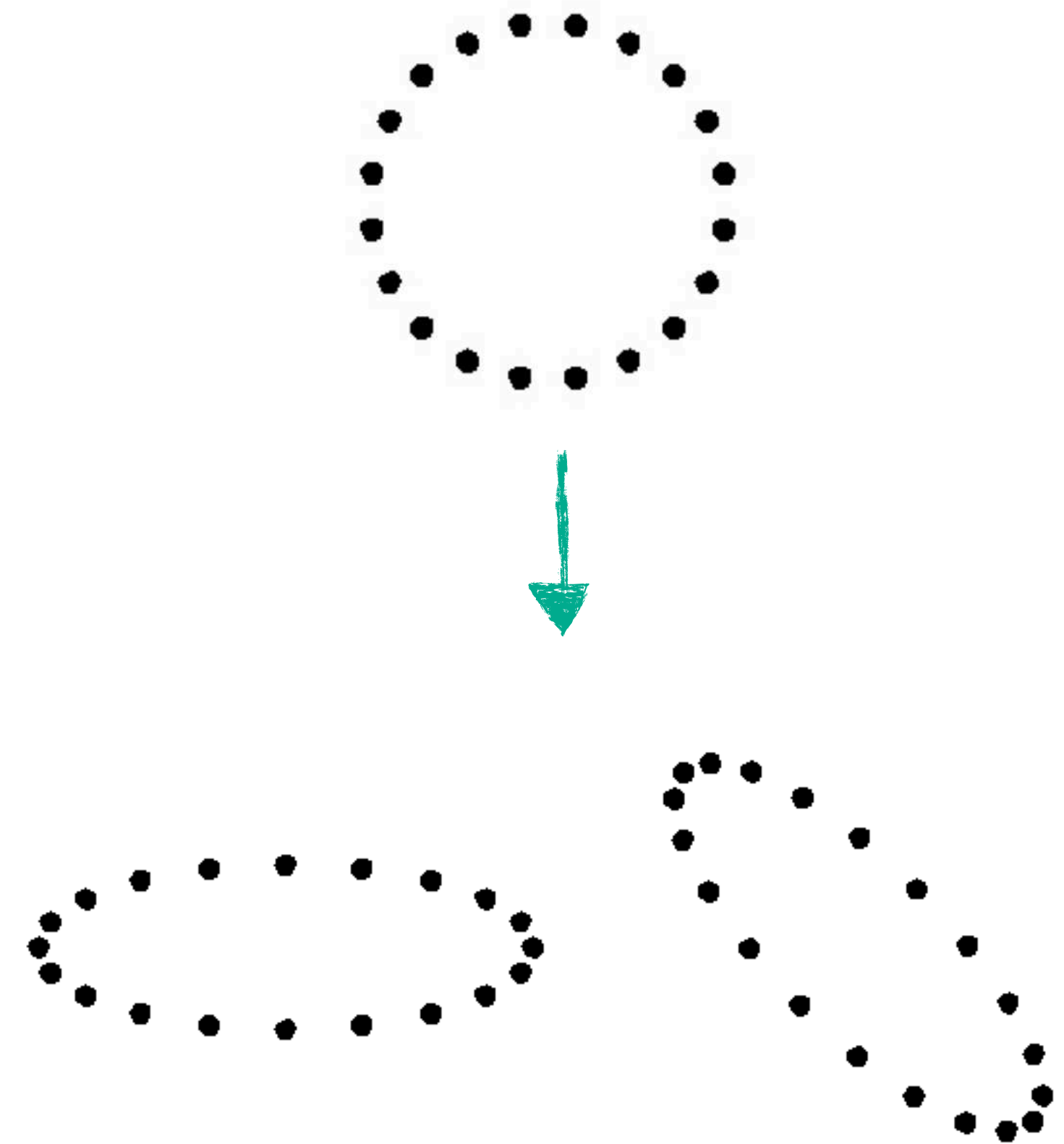


Gravitational waves

$$h \sim \frac{\Delta L}{L}$$

In General Relativity, two polarisations

$$h_+, h_\times$$



Steps to implement GWs in PHANTOM



Price et al. 2018:

- low-memory, fast, high-efficient
- widely tested for accretion physics
- we love PHANTOM

Liptai and Price 2019:

- able to capture relativistic shocks
- precise and accurate treatment of orbital dynamics
- ability to work in Kerr geometry



Steps to implement GWs in PHANTOM

Toscani, Lodato, Nealon 2019
- in post process

Steps to implement GWs in PHANTOM

Toscani, Lodato, Nealon 2019
- in post process

$$h^{\text{TT}}(t, \mathbf{n}) \propto \ddot{M}^{\text{kl}}$$

$$M^{\text{kl}} = \frac{1}{c^2} \int d\mathbf{x} T_{00} x^{\text{k}} x^{\text{l}} \Rightarrow M^{\text{kl}} = \sum_{\text{a}} m_{\text{a}} x_{\text{a}}^{\text{k}} x_{\text{a}}^{\text{l}},$$

On the fly!

$$\ddot{M}^{\text{kl}} = \sum_{\text{a}} m_{\text{a}} (\ddot{x}^{\text{l}} x^{\text{k}} + 2\dot{x}^{\text{k}} \dot{x}^{\text{l}} + x^{\text{l}} \ddot{x}^{\text{k}})_{\text{a}}$$

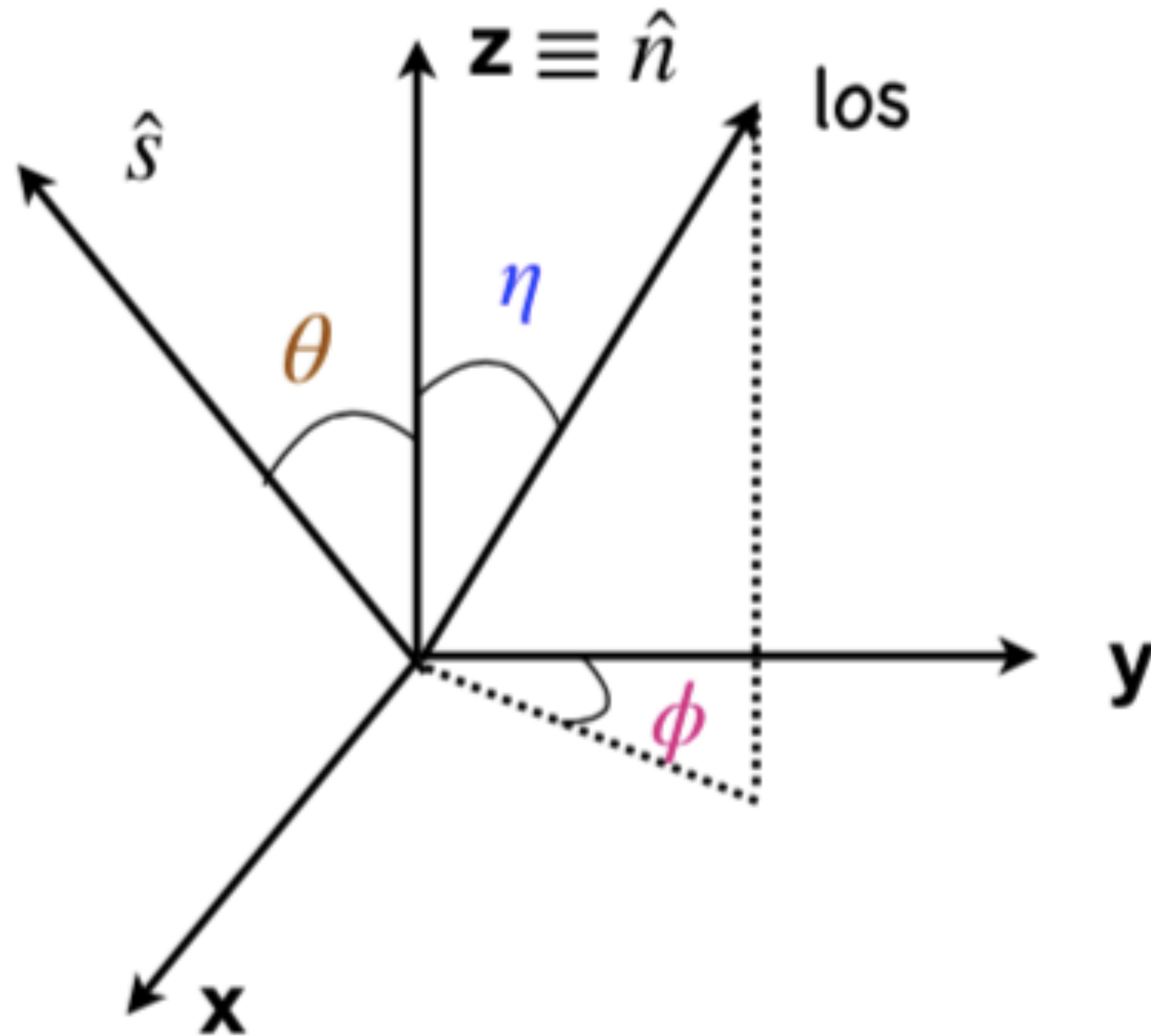
M: inertia moment of the system

a: index that runs over the number of particles

$$h_{+} h_{\times}$$

Toscani, Lodato, Price, Liptai 2022

Steps to implement GWs in PHANTOM



η : angle between the line of sight and the normal to the orbit

θ : rotation of xy plane

ϕ : angle between the projection between line of sight and y axis

Toscani, Lodato, Price, Liptai 2022

Tidal Disruption Events (TDEs)

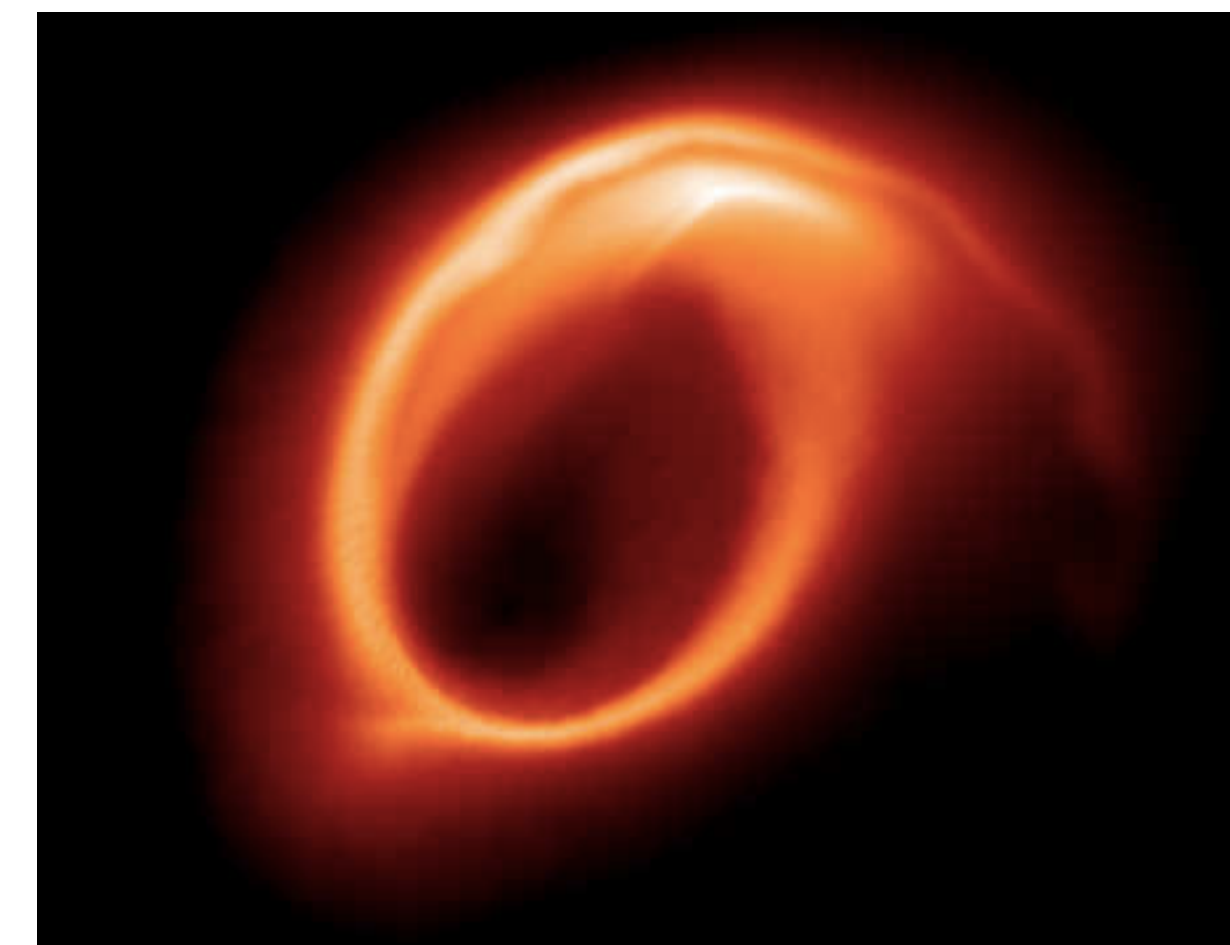
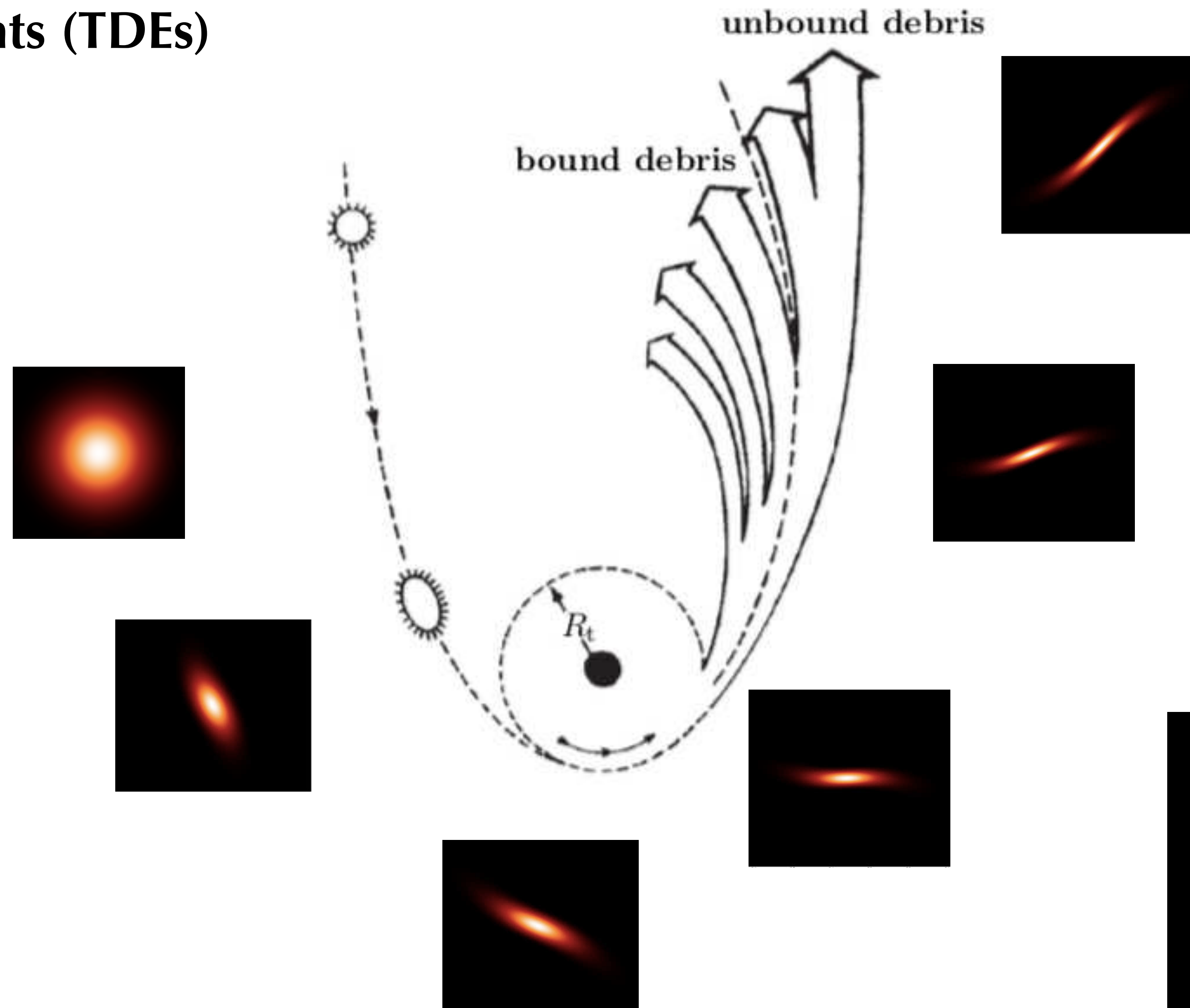


Image from Rees (1988). Snapshots by M. Toscani.
Simulation done with GRPHANTOM (Price and Liptai 2019)
and visualized with SPLASH (Price 2018).

Tidal Disruption Events (TDEs)



electromagnetic (EM) sources

accretion

100 events

optical, X-rays, radio

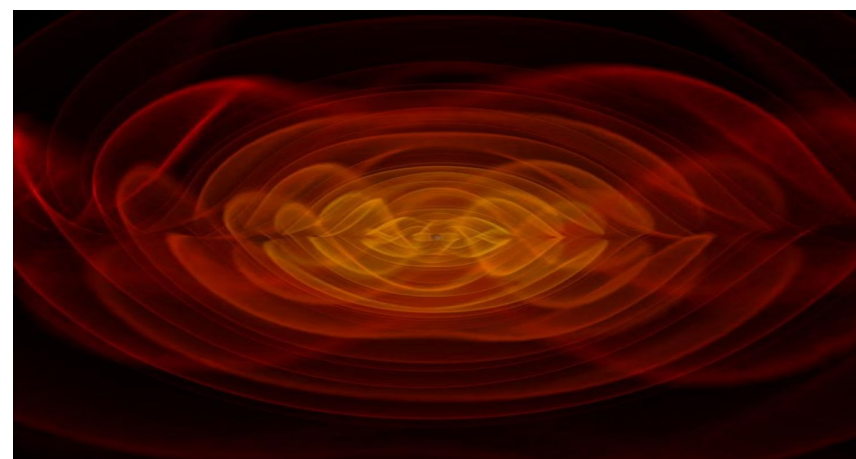
see reviews: vanVelzen et al 2020,
Saxton et al. 2020, Alexander et al. 2020



neutrino sources

few potential candidates

Stein et al. 2021, Hayasaki 2021,
Reusch et al. 2021



GW sources

disruption

Guillochon & Ramirez-Ruiz 2009,
Stone et al. 2013, Kobayashi et al.
2004, **Toscani et al. 2019,**
Toscani et al. 2021

GWs from TDEs

monochromatic GW burst

Sun-like star disrupted by a BH

$$M_h = 10^6 M_\odot$$

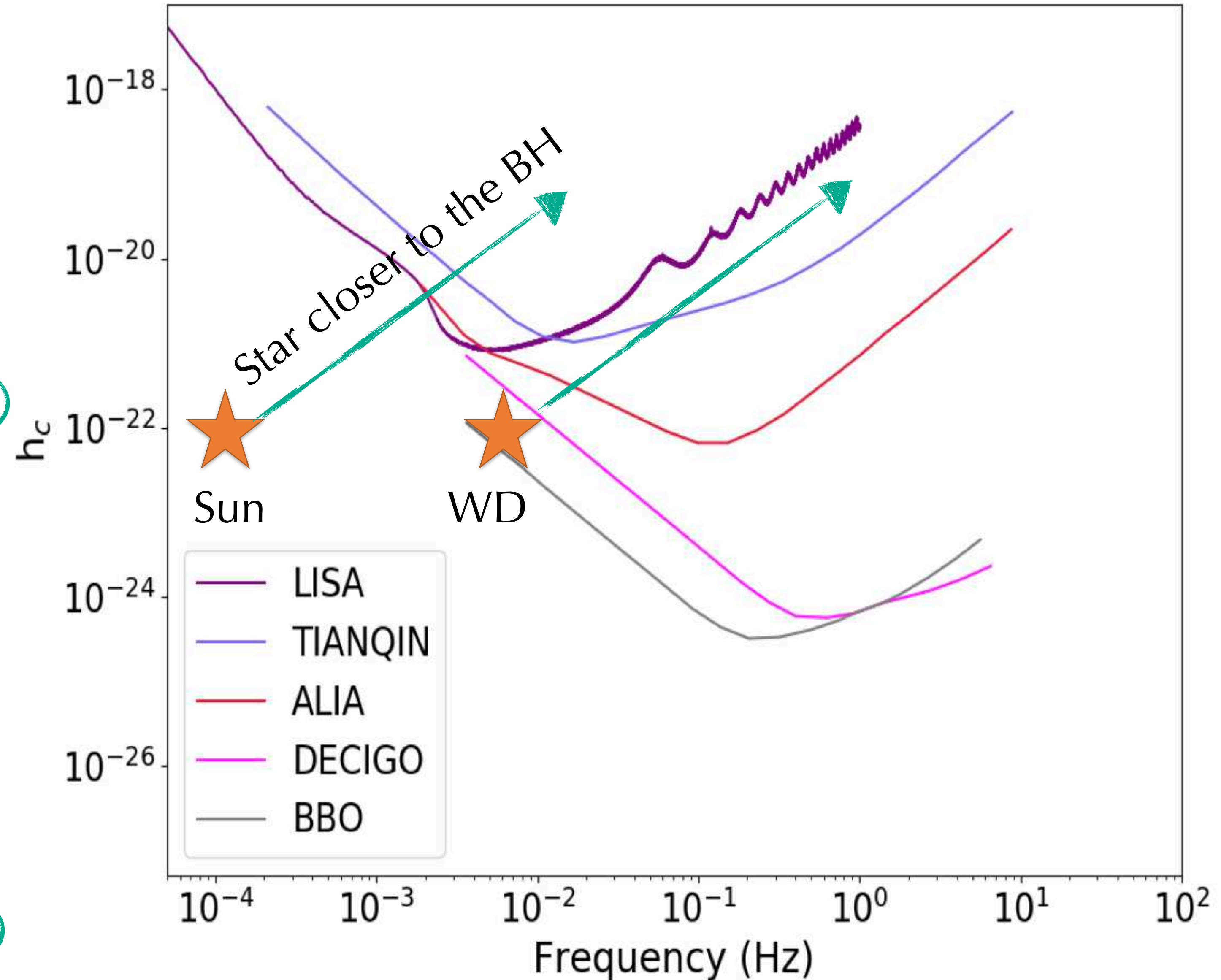
$$h \approx 10^{-22}, \quad f \approx 10^{-4} \text{ Hz}$$

White dwarf (WD)

$$M_* = 0.5 M_\odot, R_* = 0.01 R_\odot$$

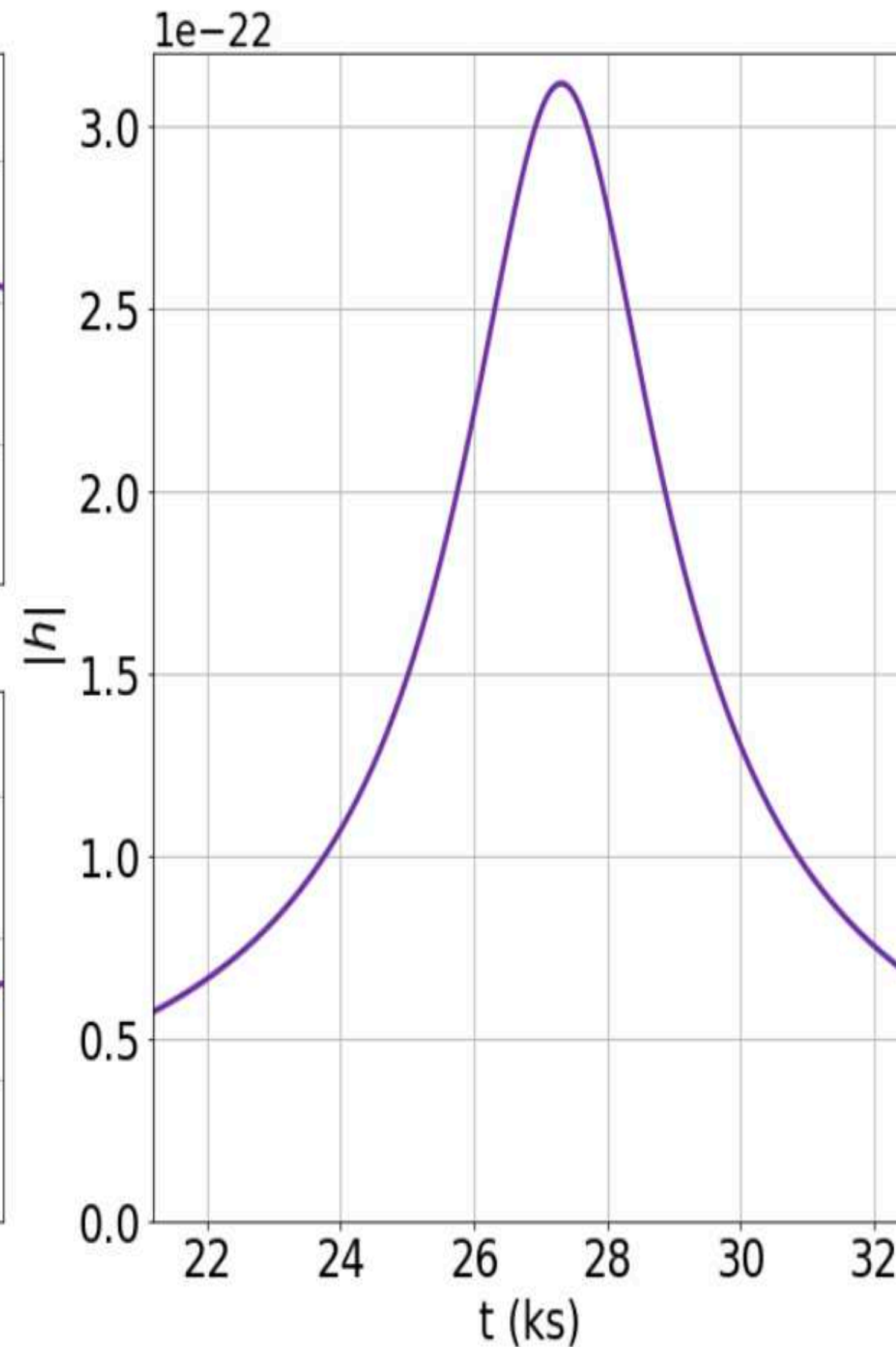
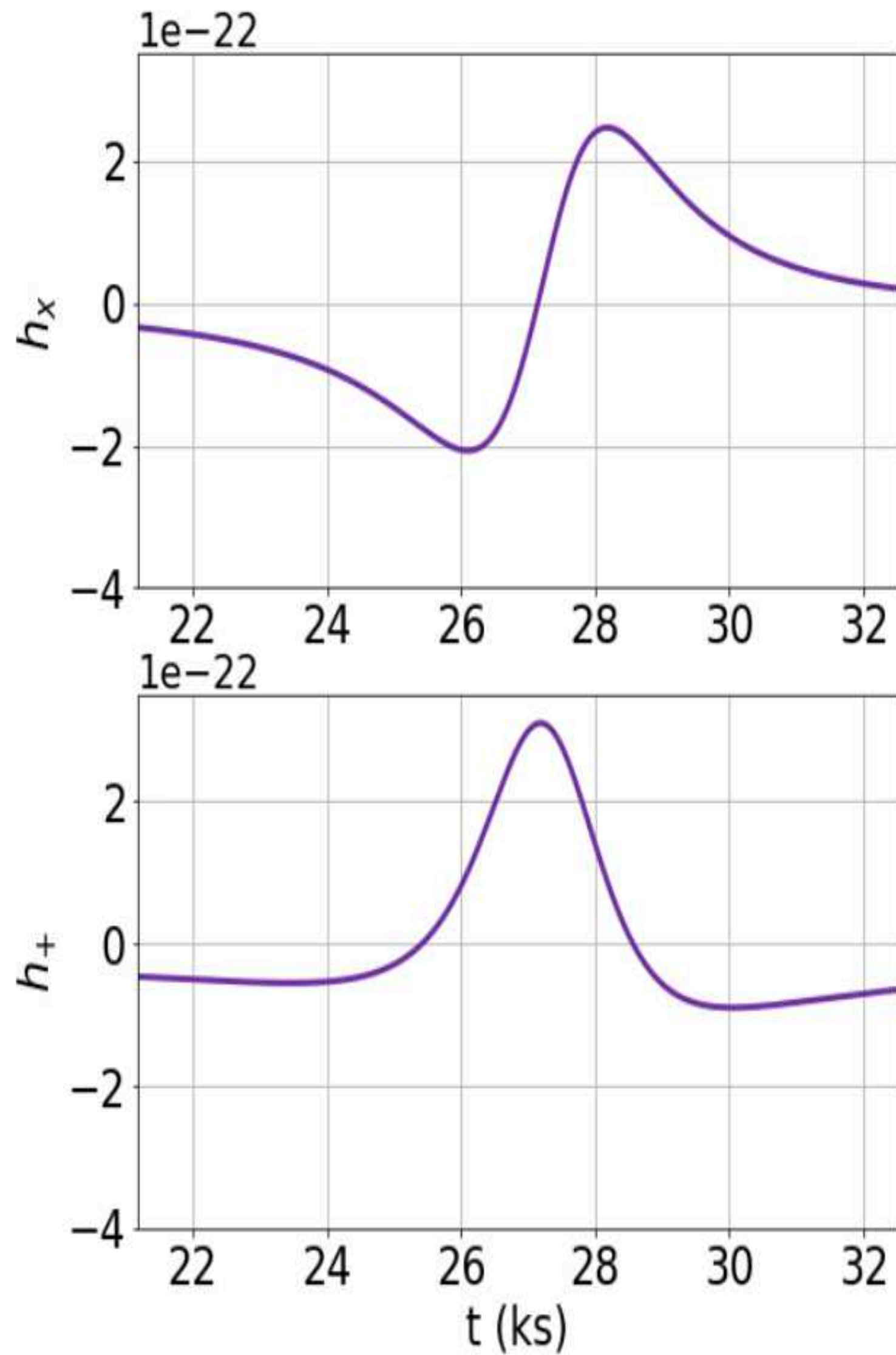
disrupted by a BH $M_h = 10^4 M_\odot$

$$h \approx 10^{-22}, \quad f \approx 10^{-2} \text{ Hz}$$



GWs from TDEs

Toscani, Lodato, Price, Liptai 2022



Standard TDE at 20 Mpc

Face-on signals

243 simulations

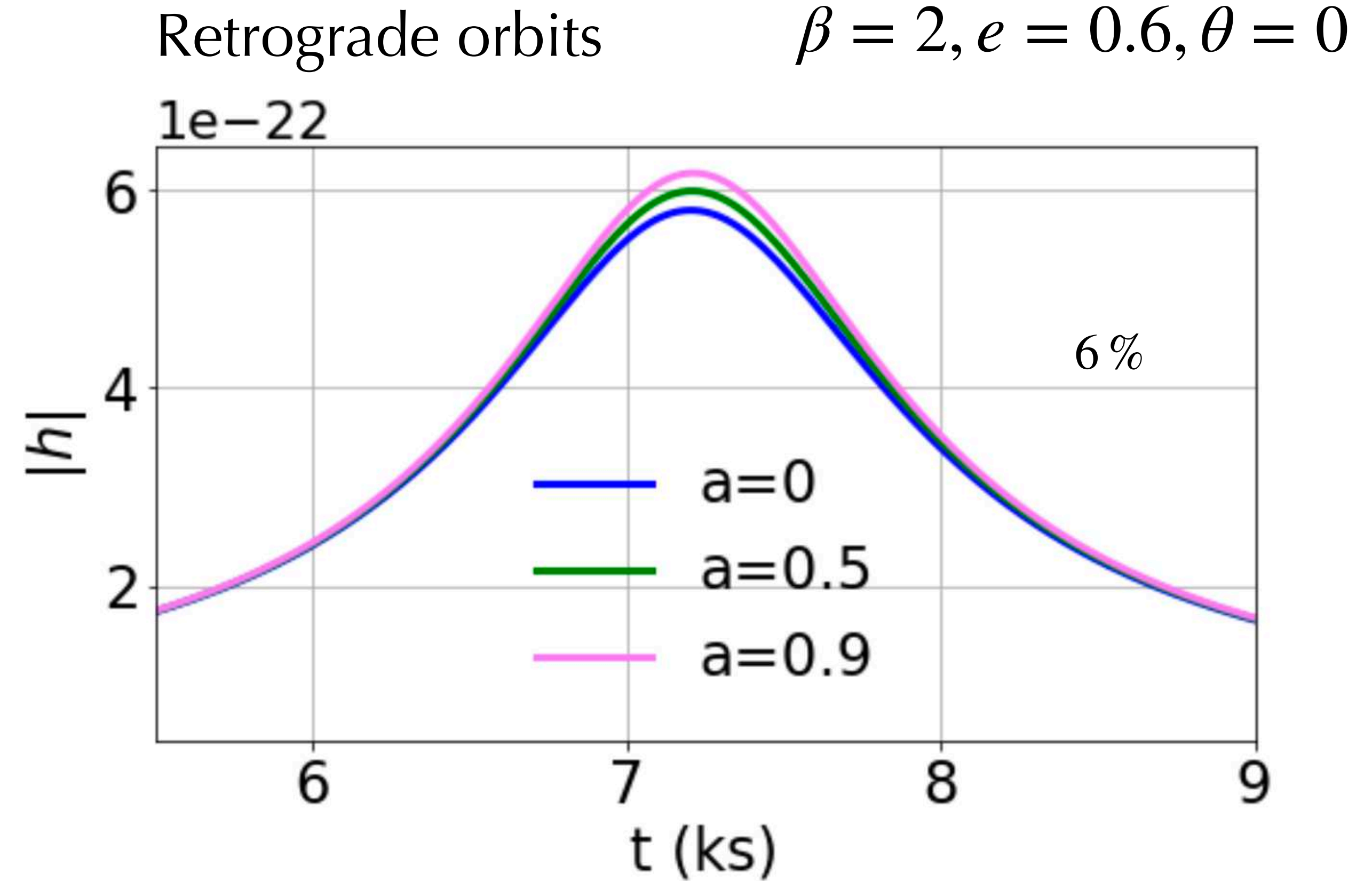
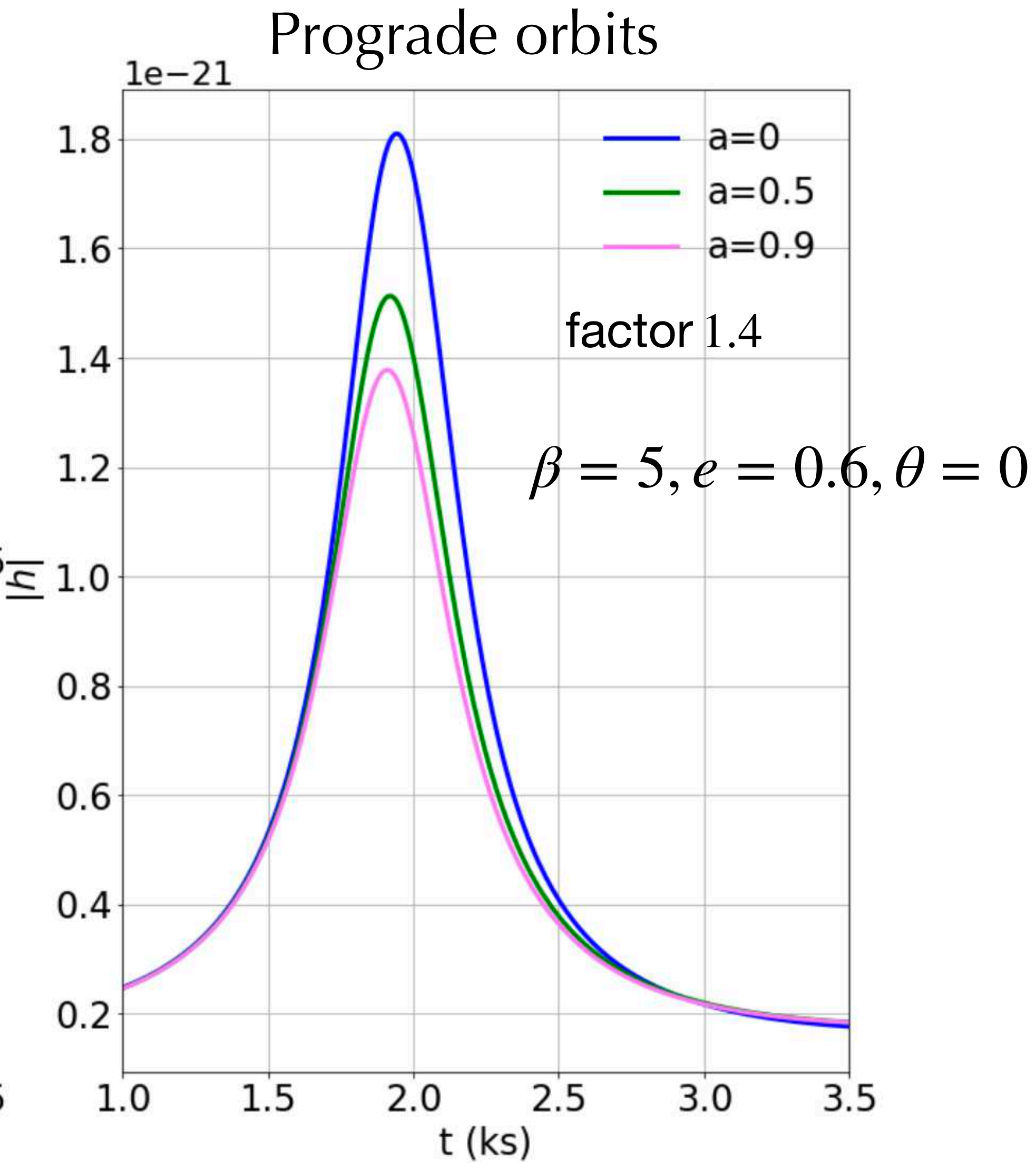
'Living' catalogue

ONLINE OPEN CATALOGUE

<https://gwcataloguetdes.fisica.unimi.it>

GWs from TDEs

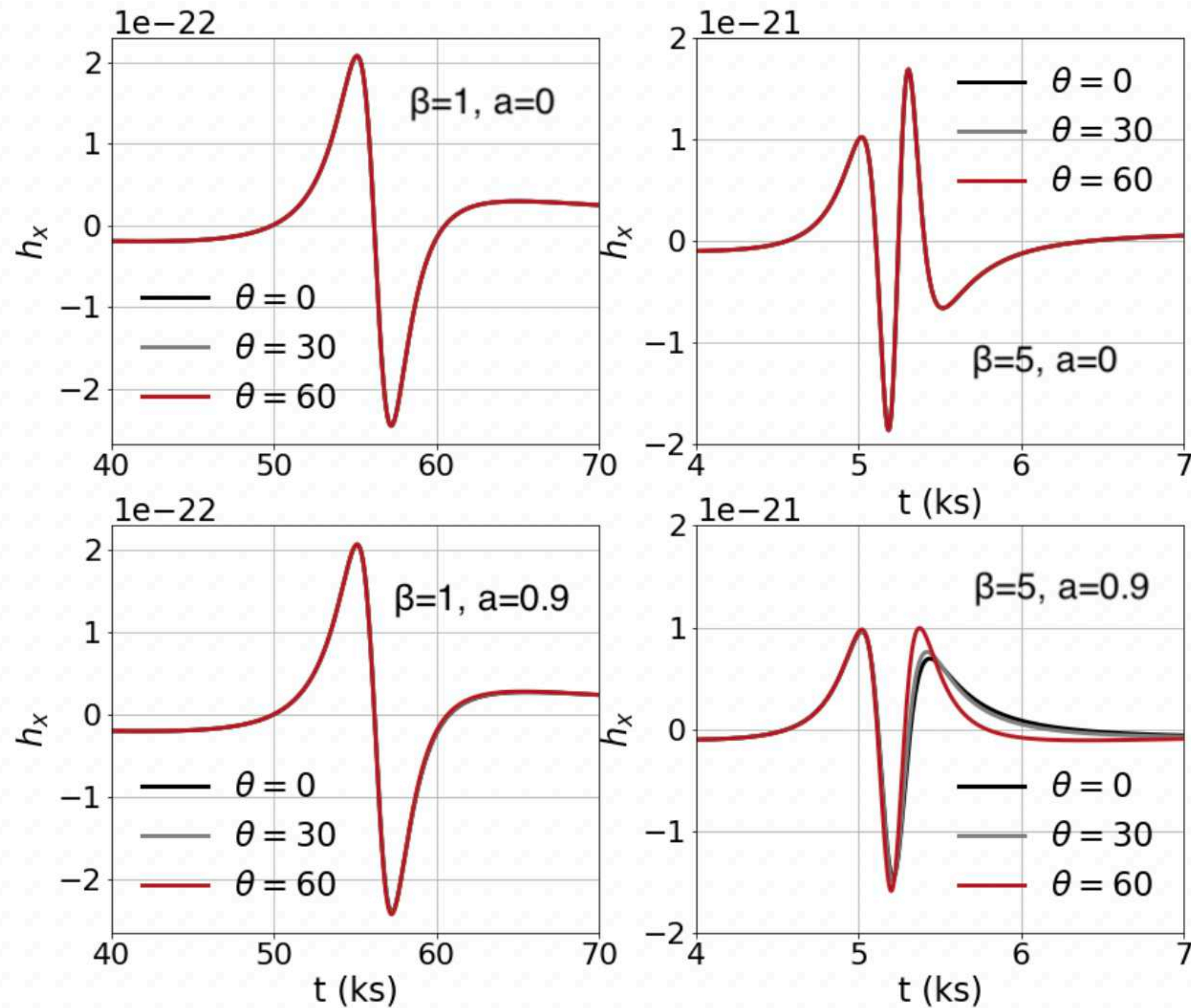
Toscani, Lodato, Price, Liptai 2022



GW signal increases for high retrograde orbits, decreases for high prograde orbits

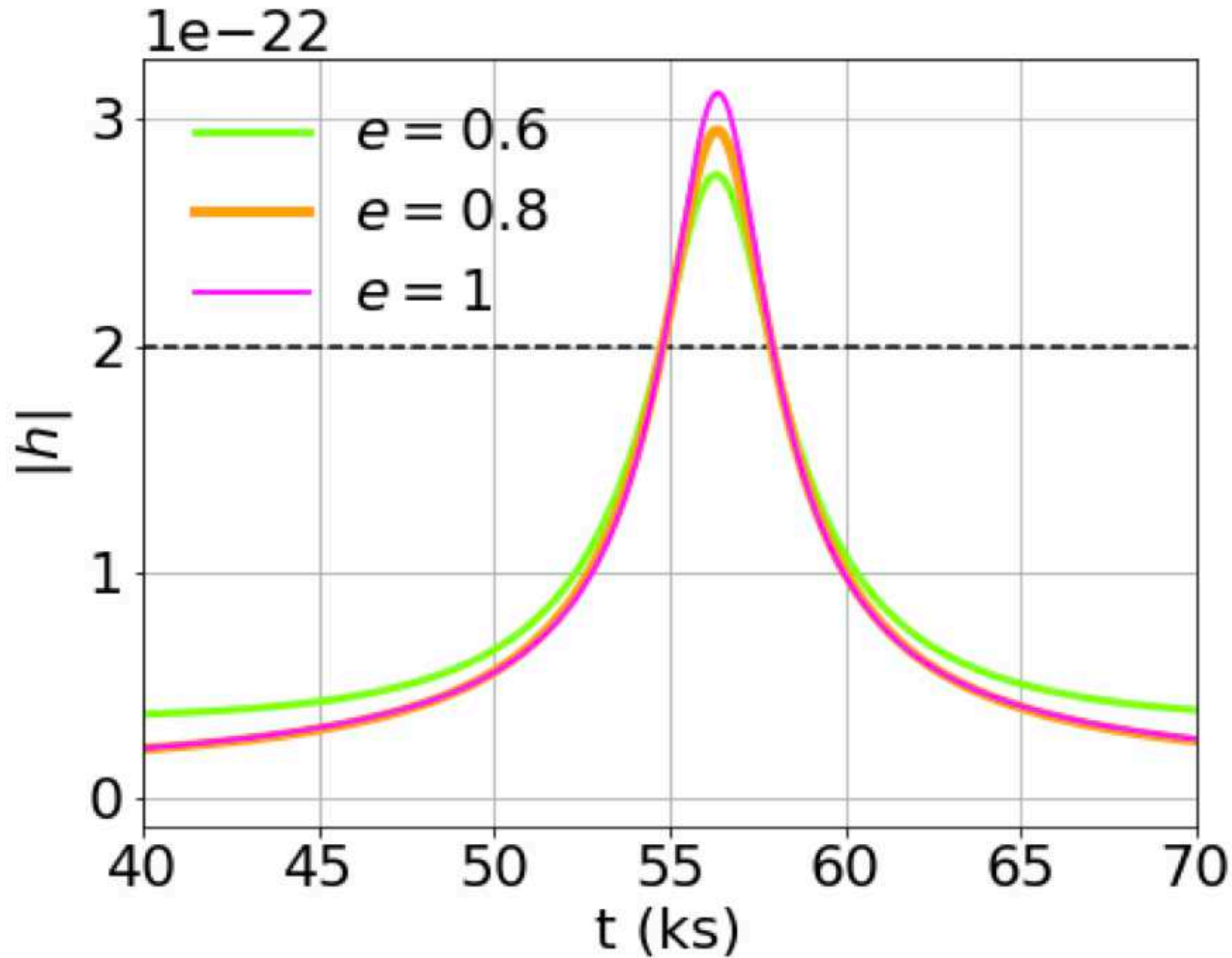
GWs from TDEs

Toscani, Lodato, Price, Liptai 2022



$$e = 0.8$$

Inclination angles affect signal only with spin AND penetrating events

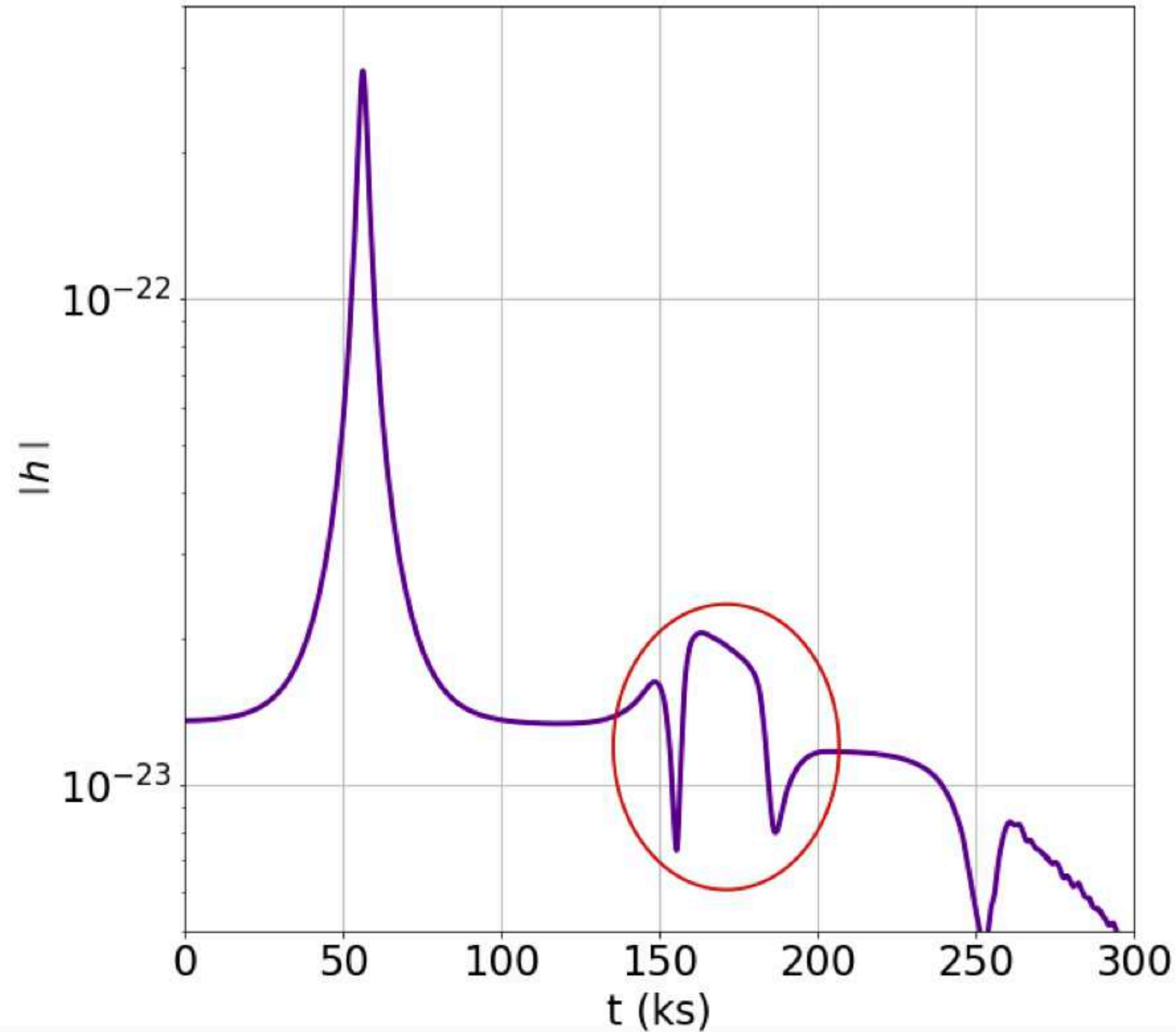


$$\beta = 1, \theta = 0$$

GW signal increases for higher eccentricities

GWs from TDEs

Toscani, Lodato, Price, Liptai 2022



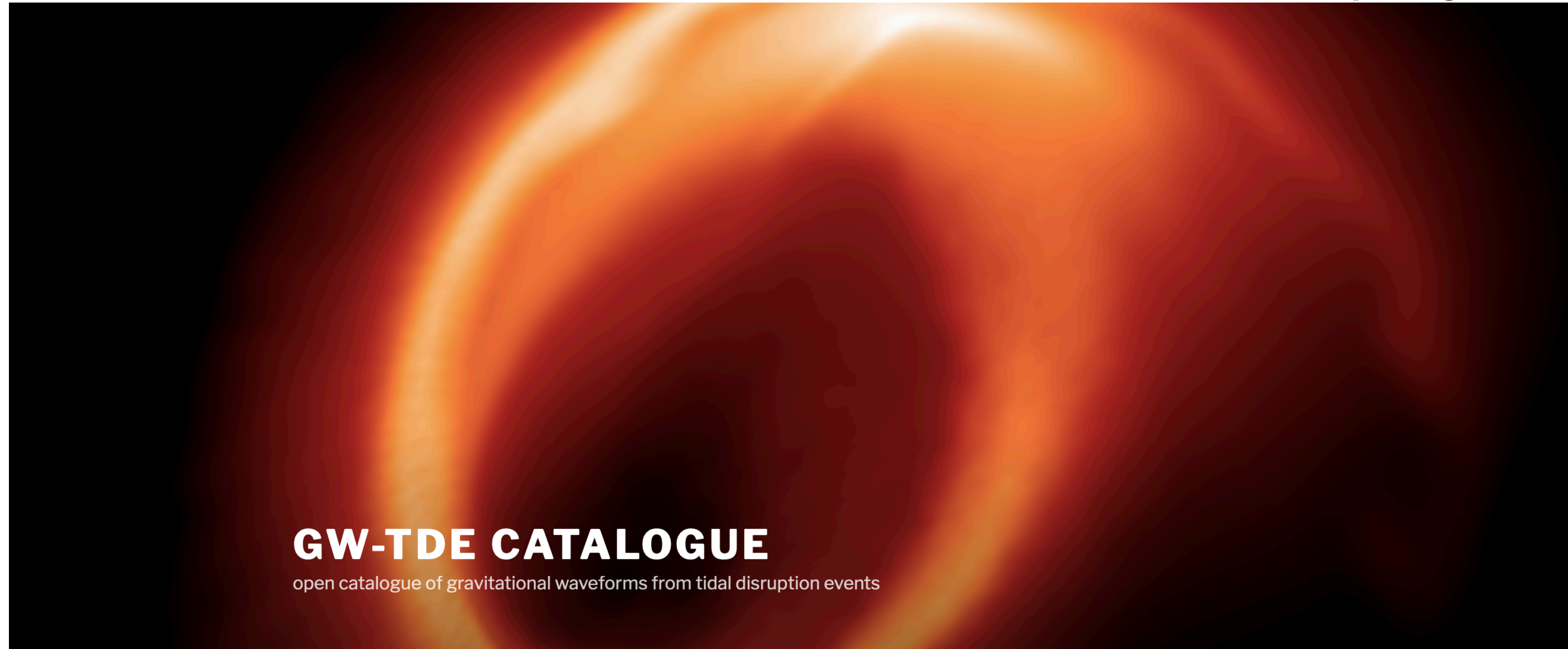
$$\beta = 1, e = 0.8$$

With interferometers sensitive enough,
combined detection of GWs and EM

What's next

1) catalogue needs to be enlarged (maintained by myself)

<https://gwcataloguetdes.fisica.unimi.it>



Home Catalogue Download Links

self-contained measurement of H_0 via Hubble's law,
completely independent of any specific cosmological models
(idea suggested by Wong 2023)

ABOUT THIS SITE

martina.toscani@l2it.in2p3.fr

PHANTOM and MCFOST Users Workshop 2023

What's next

2) Explore other systems or variations of TDEs

Partial TDE

TDE around a BH binary

Binary stars tidally disrupted

GW signal emitted by BH binary surrounded by a disc with some particular stuff going on

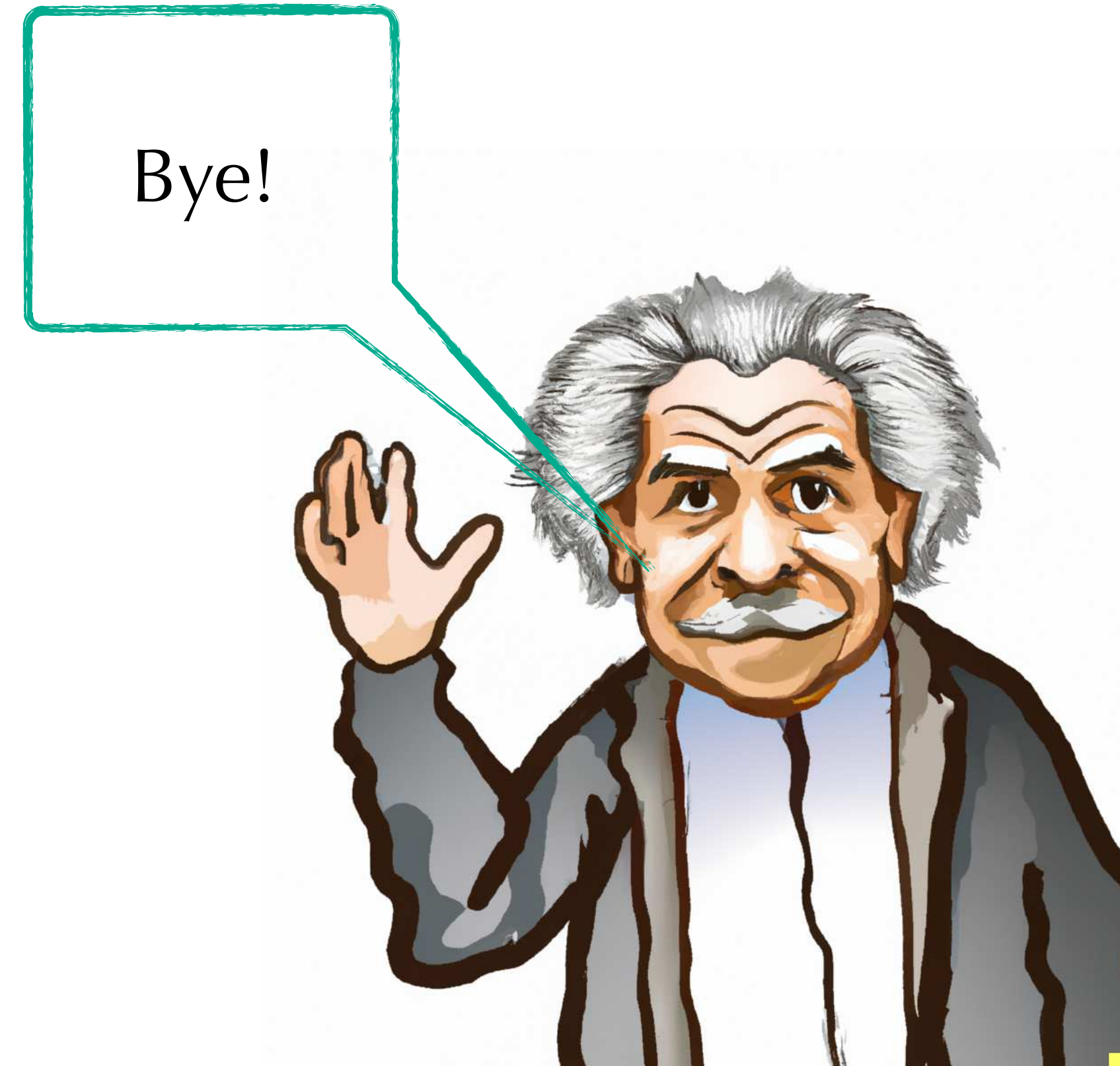
Planet disruption

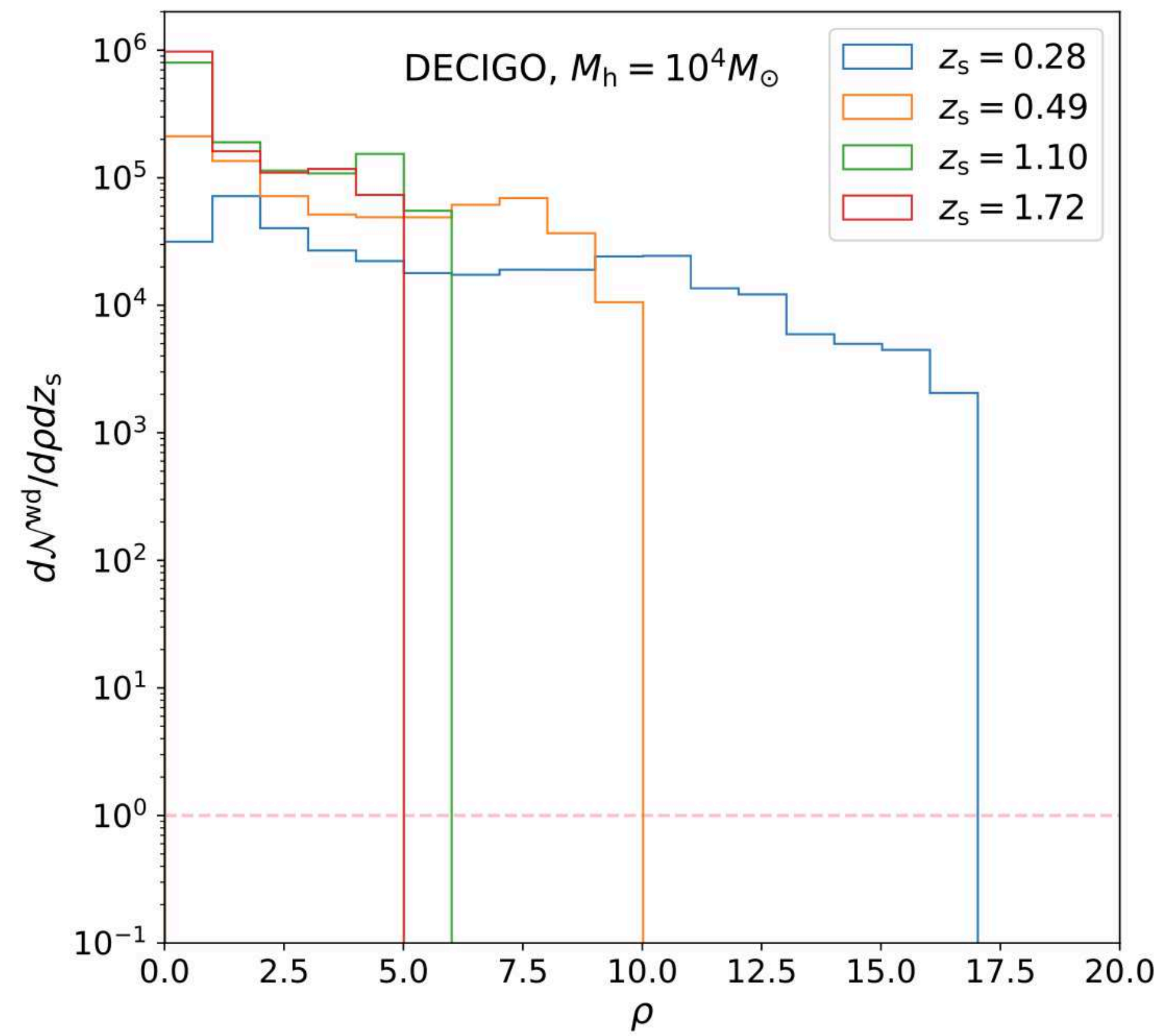
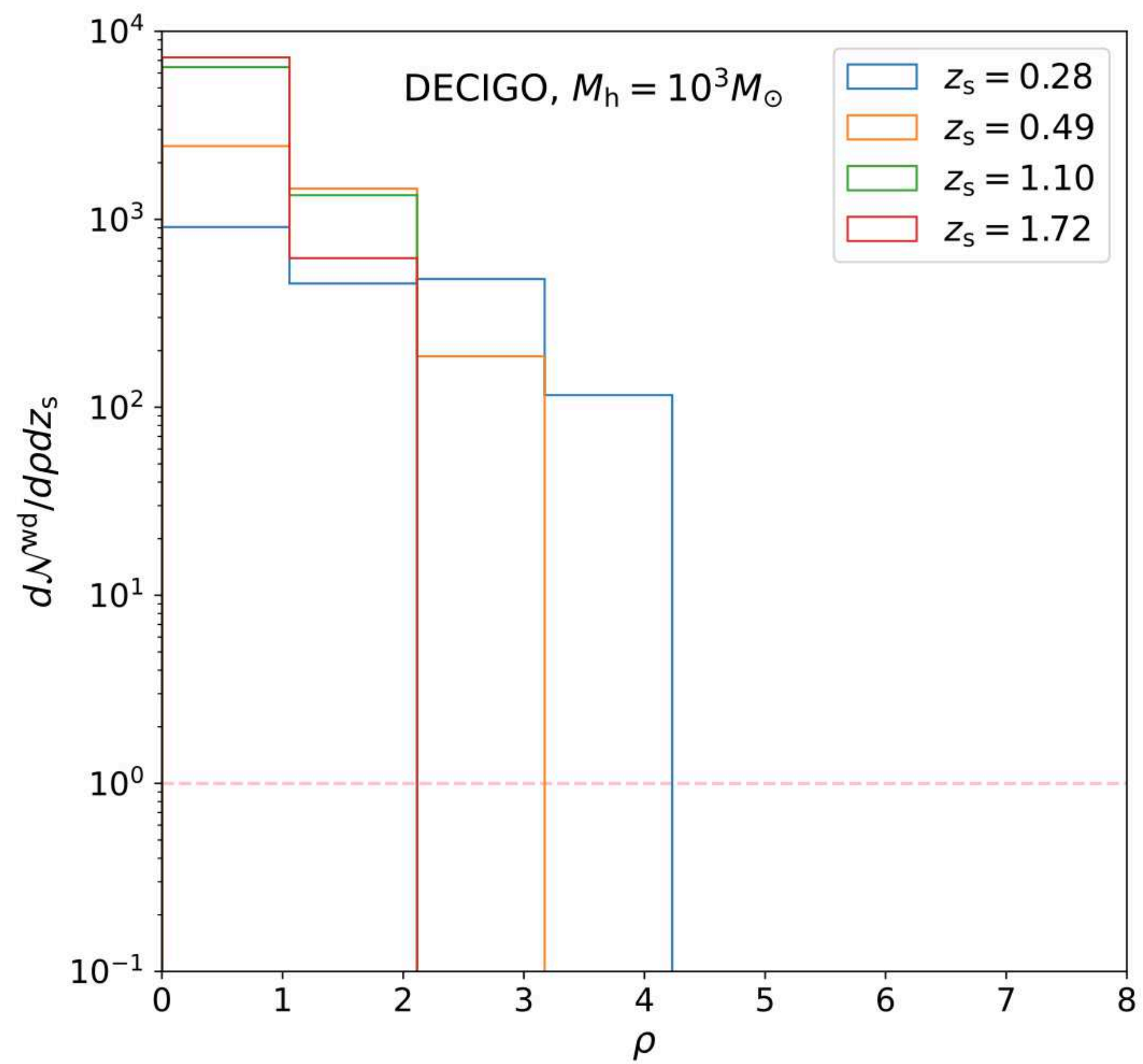
3) PostNewtonian formalism

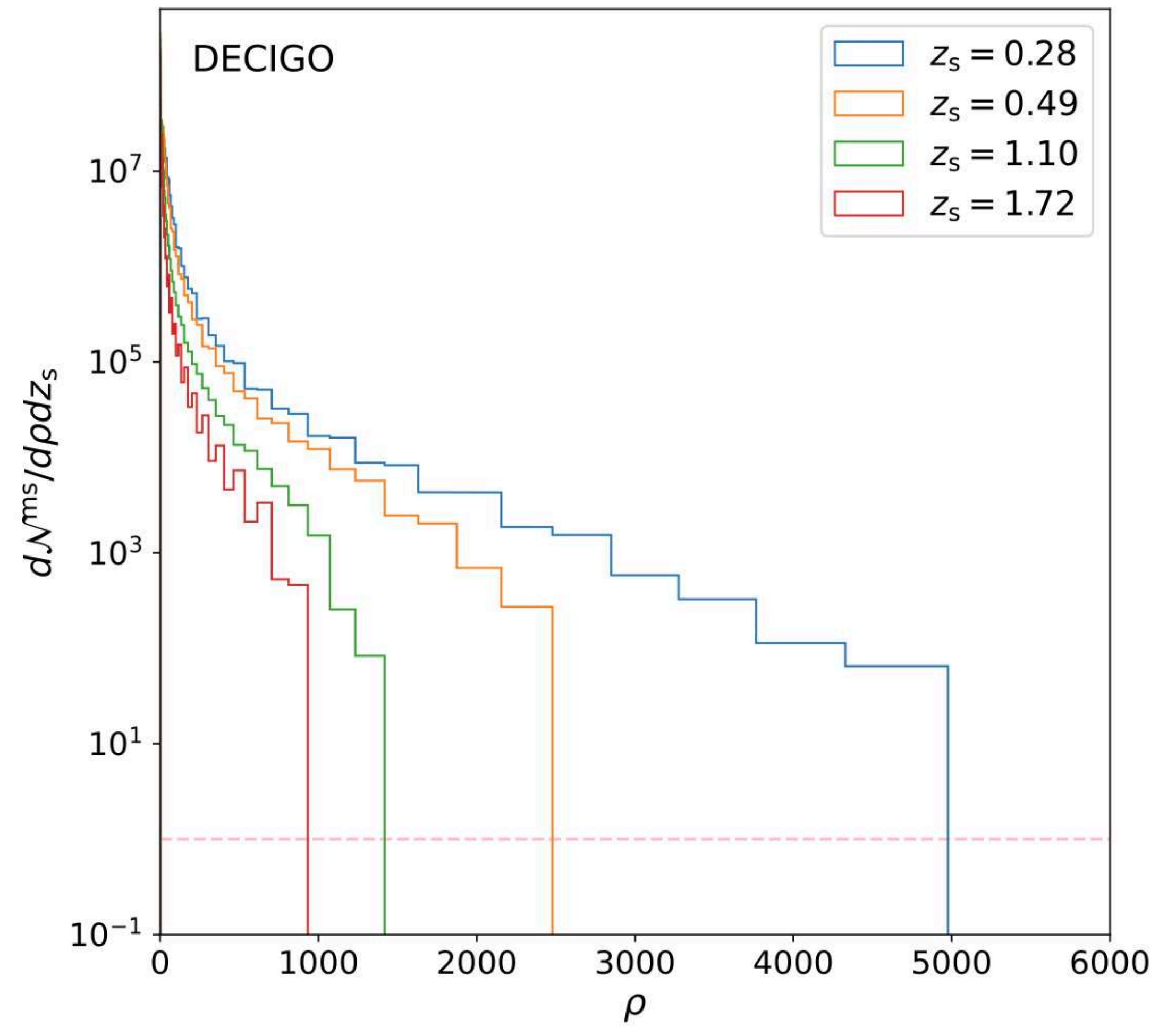
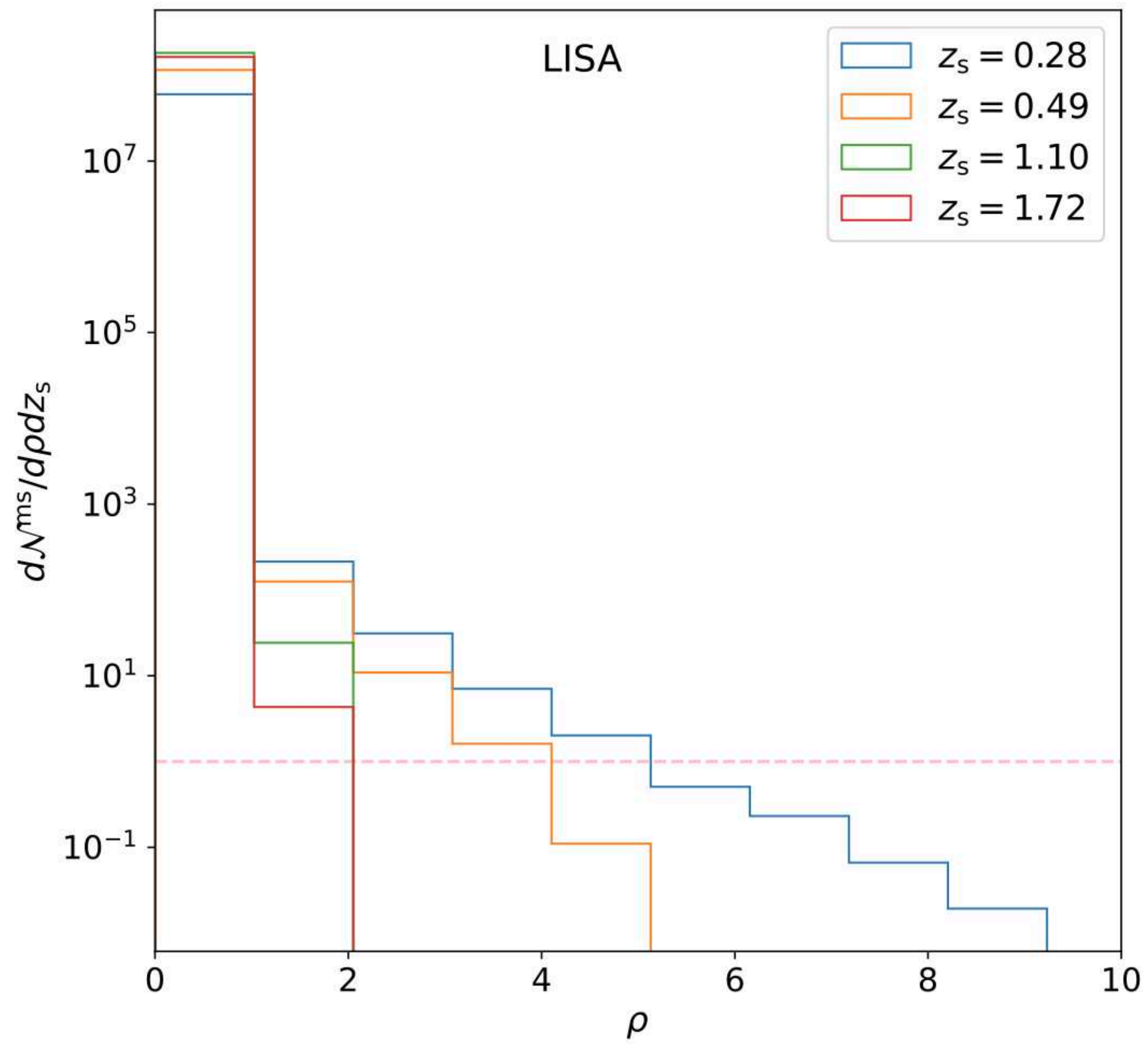
Fin

Thanks for your attention!

Questions?







$$|v_\phi| = \frac{c^3}{2\pi G M_h} \left| \frac{1}{(2r/r_g)^{3/2} \pm a} \right|$$

$$v_{\text{kepl}} = \left[\frac{GM_h}{r_p} (1 + e) \right]^{1/2} \propto (1 + e)^{1/2}$$